

# Twilight for the energy conditions?

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## Abstract:

The **energy conditions** of general relativity permit one to deduce very powerful and general theorems.

However, the **energy conditions** are now realized to be a lot less secure than they once seemed:

— There are quantum effects that violate all of the energy conditions.

— There are relatively benign classical systems that violate all the energy conditions.

This opens up a **Pandora's** box of rather disquieting possibilities.

Is it twilight time for the energy conditions?

## Overview:

**Classical:** In classical general relativity the energy conditions are used to prove lots of general theorems...

**Quantum:** Everyone who thought about it expected the energy conditions to eventually break down once one reaches the **Planck** slop...

**Surprise 1:** The energy conditions already fail miserably in **semi-classical quantum gravity**.

Failures occur at first order in  $\hbar$ , long before one reaches the **Planck** slop...

The quantum failures are widespread, albeit small...

**Surprise 2:** The energy conditions also fail miserably in quite reasonable **classical systems**.

## Energy conditions:

### Standard (pointwise) energy conditions:

**TEC** — trace energy condition  
(now abandoned).

**NEC** — null energy condition.

**WEC** — weak energy condition.

**SEC** — strong energy condition.  
(*aka* unphysical energy condition.)

**DEC** — dominant energy condition.

### Standard averaged energy conditions:

**ANEC** — averaged null energy condition.

**AWEC** — averaged weak energy condition.

**ASEC** — averaged strong energy condition.

### Standard propaganda:

Physically reasonable Lagrangians give classical theories satisfying the energy conditions...

The standard propaganda is **wrong**.

## Definition:

In a **Lorentzian** spacetime:

The null energy condition [NEC] is said to hold at a point  $x$ , if for all null vectors  $k$

$$T_{\mu\nu} k^\mu k^\nu \geq 0.$$

NEC Is the weakest pointwise energy condition in common use.

## Notes:

DEC  $\Rightarrow$  WEC  $\Rightarrow$  NEC.

SEC  $\Rightarrow$  NEC.

SEC does *not* imply WEC —  
— the terminology is misleading.

## Uses of the Energy Conditions:

- **Penrose** singularity theorem (WEC).
- **Hawking–Penrose** singularity theorem (SEC).  
[Relevant to the cosmological singularity.]
- **Tipler's** version of the **Hawking–Penrose** singularity theorem (WEC+ASEC).
- **Schoen–Yau** positive mass theorem (DEC).
- **Witten's** variant positive mass theorem (DEC).

Hypotheses are all **stronger** than the NEC.

## Uses of the ANEC:

ANEC is the weakest averaged energy condition in common use.

The ANEC is used as input hypotheses in proving:

- Focussing theorems for null geodesics.  
[Borde]
- Generalized Penrose singularity theorem.  
[Roman]
- Topological censorship theorem.  
[Friedman–Schleich–Witt]
- Generalized positive mass theorem.  
[Penrose–Sorkin–Woolgar]

## Quantum violations of the energy conditions:

- 2-particle **Fock** states. (NEC+)
- **Casimir** vacuum. (NEC+)  
[**DeWitt**, **Einstein** Centenary Survey]
- **Hawking** radiation. (NEC+)
- Squeezed vacuum. (WEC+DEC)  
[**Morris–Thorne**]
- Conformal anomaly (NEC+)  
[**Visser**, PLB 349 (1995) 443–447;  
gr-qc/9409043].
- Gravitational vacuum polarization (NEC+)  
[**Visser**, gr-qc/9604007; 9604008;  
9604009; 9703001].
- Cosmological particle production. (SEC)



## Classical violations of the energy conditions:

### “Observational”

- Cosmological inflation. (SEC)  
[Minimally coupled massive scalar]
- Cosmological inflation. (NEC+)  
[Conformally coupled massive scalar]
- Galaxy formation:  $0 < z < 10$ . (SEC)  
[Visser, Science 276 (4 April 1997) 88]
- Accelerating universe. (SEC)

## Classical violations of the energy conditions:

### “Theoretical”

- Tolman wormholes (SEC)  
[Hochberg, Molina–París, Visser]
- Massless conformally coupled scalar (ANEC+)  
QFT: new improved energy-momentum  
[Barceló–Visser, gr-qc/9908029, PLB]  
[Barceló–Visser, gr-qc/0001099, Cosmo99]
- Non-minimally coupled scalar (ANEC+)  
massive/massless  
[Wald–Flanagan, gr-qc/9602052, PRD]  
[Barceló–Visser, gr-qc/0003025, CQG]
- String moduli fields (ANEC+)  
[Barceló–Visser, gr-qc/0001099, Cosmo99]
- Negative tension branes (ANEC+)  
[Barceló–Visser, hep-th/0004022, NPB]

## Tolman wormhole:

Q: What are the *minimal* conditions for a “bounce”?

D: (Tolman wormhole  $\equiv$  “bounce”.)

A: Perform a model-independent analysis of the geometry near a bounce, along the lines of the Morris–Thorne analysis for traversable wormholes.

Details:

gr-qc/9810023, PLB455 (1999) 90-95

[Molina-París, Visser]

gr-qc/9810029, PRD59 (1999) 044011

[Hochberg, Molina-París, Visser]

Flare-out at the bounce  $\Rightarrow$  SEC violated at or near the bounce.

## Notes:

SEC violations are a *necessary* but not *sufficient* condition for a “bounce” .

You do not *need* to violate NEC, WEC, or DEC to get a “bounce” .

If you believe inflation you have already abandoned the SEC anyway.

Inflation will not *guarantee* a bounce, but it opens the door.

## New improved stress tensor plus gravity:

Take a massless conformally coupled scalar field and add **Einstein** gravity.  
(Static, spherically symmetric.)

Absurdly easy problem; surprising result.

— NEC and ANEC are often violated.

— There is a three-parameter class of exact solutions (total mass, scalar charge, scalar field at infinity).

Special cases:

(1) **Schwarzschild/ anti-Schwarzschild.**

(2) **Naked singularities.**

(3) **Naked singularities hiding behind wormhole throats.** (Not really “traversable”.)

(4) **Traversable wormholes with two asymptotically flat regions.**

Details: gr-qc/9908029, **Barceló–Visser**, PLB.

## Non-minimally coupled scalars plus gravity:

Take a generic non-minimally coupled scalar field and add **Einstein** gravity.  
(Static, spherically symmetric.)

Absurdly easy problem; surprising result.

— NEC and ANEC are often violated.

— There is a four-parameter class of exact solutions (total mass, scalar charge, scalar field at infinity, and curvature coupling).

Special cases:

- (1) **Schwarzschild/ anti-Schwarzschild.**
- (2) **Naked singularities.**
- (3) **Naked singularities hiding behind wormhole throats.** (Not really “traversable”.)
- (4) **Traversable wormholes with two asymptotically flat regions.**

For  $\xi \leq 0$  no traversable wormholes.

For  $\xi > 0$  get traversable wormholes.

Details: gr-qc/0003025, **Barceló–Visser**, CQG.

## Negative tension branes plus gravity:

Take a negative tension brane and add Einstein gravity.

(Static, spherically symmetric, (3+1)-D.)

(Negative tension branes are now extremely common in brane cosmology, and variants of the Randall–Sundrum scenario. More than 50 papers as of July 2000.)

Absurdly easy problem, surprising result.

— NEC and ANEC are always violated.

— There is a four-parameter class of exact solutions (total mass, brane tension, bulk cosmological constant, electric charge).

Special cases:

Traversable wormholes with two asymptotically de Sitter regions.

Details: hep-th/0004022, Barceló–Visser, NPB.

## Specific Implications:

Don't focus on the specific technical details. (Naked singularities, traversable wormholes). The main points are:

- Classical violations of the NEC arise in these very reasonable classical systems.
- Conformally coupled scalars are from a QFT perspective the preferred choice — corresponding to the new improved stress-energy tensor.
- String moduli are generic.
- Negative tension branes are ubiquitous.
- It's the fact that you get classical NEC violations in such simple physical systems that's worrying — the fact that these NEC violations are big enough to support traversable wormholes is a bonus.



## General Implications:

- We do not currently have an acceptable **positive mass theorem**.
- We do not currently have an acceptable **singularity theorem**.
- We do not currently have an acceptable **topological censorship theorem**.
- **Traversable wormholes** almost begin to look physically reasonable.
- **Tolman wormholes** almost begin to look physically reasonable.
- This opens up a whole mess of **weird** possibilities...

## Conclusions:

We need:

- Improved understanding of just what conditions can sensibly be put on the stress-energy tensor.  
(Quantum inequalities?)
- Improved energy conditions — you do not want a free-for-all.  
(Arbitrary stress-energy tensor  $\Rightarrow$  arbitrarily weird physics.)
- Improved (positive-mass/ singularity/ censorship) theorems of all types.

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